Gaussian quantum Monte-Carlo Methods phase-space simulations of many-body systems

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Overview

- introduction to phase-space representations
 - +P method
- Gaussian operator bases
 - overcompleteness and differential properties
- quantum evolution in real or imaginary time
 - mapping to stochastic phase-space equations
- application to Hubbard model
 - improvements through symmetry and gauge considerations

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Wigner distribution

- A classical state can be represented by a joint probability distribution in phase space $P(\mathbf{x}, \mathbf{p})$
- 1932: Wigner¹ constructed analogous quantity for quantum states:

$$W(x,p) = \frac{2}{\pi} \int dy \psi^* (x-y) \psi(x+y) \exp\left(-\frac{2iyp}{\hbar}\right)$$

✓ Wigner function gives correct marginals: $\frac{\int dx W(x,p) = 2\hbar P(p)}{\int dp W(x,p) = 2\hbar P(x)}$

 \checkmark but it is not always positive \rightarrow not a true joint probability

1. E. P. Wigner, Phys. Rev. 40, 749 (1932).

P and **Q** distrubutions

- other phase-space distributions were also defined by
 - Husimi¹:

$$Q(\alpha) = \frac{1}{\pi} |\langle \Psi | \alpha \rangle|^2$$

and

• Glauber² and Sudarshan³:

$$|\Psi\rangle\langle\Psi|=\int d^2\alpha P(\alpha)|\alpha\rangle\langle\alpha|$$

- 1. K. Husimi, Proc. Phys. Math. Soc. Jpn. 22, 264 (1940).
- 2. R. J. Glauber, Phys. Rev. 131, 2766 (1963).
- 3. E. C. G. Sudarshan, Phys. Rev. Lett. 10, 277 (1963).

Phase-space distributions

• Wigner, Q and P all defined in terms of coherent states

$$\widehat{a}|\alpha\rangle = \alpha|\alpha\rangle; \ |\alpha\rangle \equiv \widehat{D}(\alpha)|0\rangle \equiv \exp[\alpha\widehat{a}^{\dagger} - \alpha^{*}\widehat{a}]|0\rangle$$

- the distributions are interrelated by Gaussian convolutions
- correspond to different choices of orderings:

$$\langle \hat{a}^{\dagger} \hat{a} \rangle = \int d^2 \alpha \left(|\alpha|^2 + n \right) p(\alpha); \ n = \begin{cases} 0 & p = P \\ -\frac{1}{2} & p = W \\ -1 & p = Q \end{cases}$$



Probability distributions

to be a probabilistic representation, these functions must:

	P	W	Q
exist and be nonsingular	×	~	~
always be positive	×	X	~
evolve via drift and diffusion	×	X	X



Reversibility

- classical random process is irreversible
 - → outward (positive) diffusion
- quantum mechanics is reversible
 - → phase-space functions generally don't have positive diffusion

A solution!

- dimension doubling
 - diffusion into 'imaginary' dimensions
 - observables evolve reversibly
 - also fixes up existence and positivity

positive-P representation

- generalisation of the Glauber P by Drummond et al^{1,2}
- expand $\hat{\rho}$ over off-diagonal coherent-state projectors:

$$\widehat{\rho} = \int d^2 \alpha \, d^2 \beta \, P(\alpha, \beta) \frac{\left|\alpha\right\rangle \left<\beta\right|}{\left<\beta\right|\alpha\right>}$$

- off-diagonal coherent projectors are a very overcomplete basis
 - many equivalent $P(\alpha, \beta)$ functions for a given quantum state
 - can always find a positive one
 - time evolution of $\hat{\rho}$ maps to diffusive evolution of $P(\alpha, \beta)$
- 1. S. Chaturvedi, P. D. Drummond, and D. F. Walls, J. Phys. A 10, L187-192 (1977)
- 2. P. D. Drummond and C. W. Gardiner, J. Phys. A 13, 2353 (1980).

positive-*P* Simulations

maps state evolution onto two independent stochastic amplitudes:

$$\widehat{a} \rightarrow \alpha \ \widehat{a}^{\dagger} \rightarrow \beta^{*}$$

stochastic averages correspond to normally ordered correlations

$$\langle : f(\widehat{a}^{\dagger}, \widehat{a}) : \rangle = \int d^2 \alpha \, d^2 \alpha^+ f(\beta^*, \alpha) P(\alpha, \beta)$$

- many applications in quantum optics and ultracold atoms
- both real time (dynamics) and imaginary time (finite temperature) calculations

positive-P Applications

quantum optics

- superfluorescence:
 - **F**. Haake et al, Phys. Rev. Lett. **42**, 1740 (1979)
 - P. D. Drummond and J. H. Eberly, Phys. Rev. A 25, 3446 (1982).
- parametric amplifiers:
 - C. W. Gardiner, *Quantum Noise*, (Springer-Verlag, Berlin, 1991).
- quantum solitons:
 - S. J. Carter et al, Phys. Rev. Lett. 58, 1841 (1987)
 - P. D. Drummond, R. M. Shelby, S. R. Friberg and Y. Yamamoto, Nature **365**, 307 (1993)

ultracold gases

- ø dynamics of Bose-Einstein condensates:
 - M. J. Steel, M. K. Olsen, L. I. Plimak *et al*, Phys. Rev. **A58**, 4824 (1998)
 - P. D. Drummond, J. F. Corney, Phys. Rev. A 60, R2661-R2664 (1999)
 - K. V. Kheruntsyan, M. K. Olsen, and P. D. Drummond, Phys. Rev. Lett. 95, 150405 (2005).
- finite-temperature correlations in Bose gases:
 - P. D. Drummond, P. Deuar and K. V. Kheruntsyan, Phys. Rev. Lett. 92, 040405 (2004).

Example: Evaporative Cooling of a BEC

$$\widehat{H} = \int \mathrm{d}^3 \mathbf{x} \widehat{\Psi}^{\dagger}(\mathbf{x}, t) \left[-\frac{\hbar^2}{2m} \nabla^2 + V(\mathbf{x}, t) + \frac{U}{2} \widehat{\Psi}^{\dagger}(\mathbf{x}, t) \widehat{\Psi}(\mathbf{x}, t) \right] \widehat{\Psi}(\mathbf{x}, t)$$

• equivalent (Itô) stochastic equations (j = 1, 2):

$$i\hbar\frac{\partial\Psi_j}{\partial t} = \left[\frac{-\hbar^2}{2m}\nabla^2 + 4\pi a_0\frac{\hbar^2}{m}\Psi_j\Psi_{3-j}^* + V(\mathbf{x},t) + \boldsymbol{\xi}_j(\mathbf{x},t)\right]\Psi_j$$

quantum noise:

$$\langle \xi_{j'}(\mathbf{x}',t')\xi_j(\mathbf{x},t)\rangle = i\frac{4\pi a_0\hbar^3}{m}\delta_{jj'}\delta(\mathbf{x}-\mathbf{x}')\delta(t-t')$$

Example: Evaporative Cooling of a BEC



- 3D calulation with 20000 atoms, 32000 modes
- start with Bose gas above T_c ; finish with narrow BEC peak

X Problems!

- **X** method pushed to the limit
- ✗ breaks down for longer times, stronger interactions

What about fermions?

- +P is defined only for bosonic systems
- could we use fermionic coherent states^{1,2}? $\hat{c}|\gamma\rangle = \gamma|\gamma\rangle$
 - like bosonic counterparts, have useful completeness properties
 - require anticommuting, Grassmann variables: $\gamma_1 \gamma_2 = -\gamma_2 \gamma_1$
- Cahill and Glauber constructed fermionic representations³
- numerical simulations implemented⁴, but do not scale well
- the Grassmann algebra introduces great computational complexity
- 1. J. L. Martin, Proc. Roy. Soc. A 251, 543 (1959)
- 2. Y. Ohnuki and T. Kashiwa, Prog. Theor. Phys. 60, 548 (1978).
- 3. K. E. Cahill and R. J. Glauber, Phys. Rev. A 59, 1538 (1999)
- 4. L. I. Plimak, M. J. Collett and M. K. Olsen, Phys. Rev. A 64, 063409 (2001)

Phase-space representation

$$\widehat{\rho} = \int P(\overrightarrow{\lambda}) \widehat{\Lambda}(\overrightarrow{\lambda}) d\overrightarrow{\lambda}$$

- $P(\vec{\lambda})$ is a probability distribution
- $\widehat{\Lambda}(\overrightarrow{\lambda})$ is a suitable operator basis
- $\overrightarrow{\lambda}$ is a generalised phase-space coordinate
- $d\overrightarrow{\lambda}$ is an integration measure
- equivalent to

$$\widehat{\rho} = E\left[\widehat{\Lambda}(\overrightarrow{\lambda})\right]$$

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Operator Bases

need basis simple enough to fit into a computer, complex enough to contain the relevant physics:



Required Properties

- to enable a phase-space representation, the operators $\widehat{\Lambda}(\overrightarrow{\lambda})$ must
 - 1. form a *complete* basis for the class of physical density matrices
 - 2. possess a *second-order differential* form for all two-body operators
 - 3. represent any state by a *positive* expansion
 - 4. be *analytic* functions of the phase-space variables
- 3 and 4 are obtained with *non-Hermitian* elements in the basis

General Gaussian operators

generalisation of density operators that describe Gaussian states

- Gaussian states can be:
 - coherent (for bosons), squeezed (eg BCS), or thermal



- or any combination of these
- characterised by first-order moments: \overline{x} , \overline{p} , $\overline{x^2}$, $\overline{p^2}$, \overline{xp}
 - all higher-order moments factorise

Gaussian Basis: General form

$$\widehat{\Lambda}(\overrightarrow{\lambda}) = \Omega \sqrt{\left|\underline{\underline{\sigma}}\right|^{\mp 1}} : \exp\left[\delta \underline{\widehat{a}}^{\dagger} \left(\underline{\underline{I}} \mp \underline{\underline{I}} - \underline{\underline{\sigma}}^{-1}\right) \delta \underline{\widehat{a}}/2\right] :$$

relative displacement: $\delta \underline{\hat{a}} = \underline{\hat{a}} - \underline{\alpha}$ annihilation and creation operators: $\underline{\hat{a}} = (\widehat{a}_1, ..., \widehat{a}_M, \widehat{a}_1^{\dagger}, ..., \widehat{a}_M^{\dagger})$ coherent offset: $\underline{\alpha} = (\alpha_1, ..., \alpha_M, \alpha_1^+, ..., \alpha_M^+)$, ($\underline{\alpha} = 0$ for fermions) covariance: $\underline{\underline{\sigma}} = \begin{bmatrix} \mathbf{n}^T \pm \mathbf{I} & \mathbf{m} \\ \mathbf{m}^+ & \mathbf{I} \pm \mathbf{n} \end{bmatrix}$, $\underline{\underline{I}} = \begin{bmatrix} \pm \mathbf{I} & \mathbf{0} \\ \mathbf{0} & \mathbf{I} \end{bmatrix}$. upper signs: bosons; lower signs: fermions

Extended phase space

$$\overrightarrow{\lambda}$$
 = ($\Omega, \alpha, \alpha^+, n, m, m^+$)

- Ω : weight (for gauges and unnormalised density matrices)
- α and α^+ : coherent amplitudes
- n: complex number correlations (normal fluctuations)
- m and m^+ : independent squeezing (anomalous) correlations

 $\implies \text{Hilbert-space dimension:} \sim 2^{M} \text{ for fermions, } \sim N^{M} \text{ for bosons}$ $\implies \text{phase-space dimension:} \leq \begin{cases} 2(1 - M + 2M^{2}) & \text{for fermions} \\ 2(1 + 3M + 2M^{2}) & \text{for bosons} \end{cases}$

Moments

• observables are weighted averages ($E[A] \equiv \int d\lambda \Omega A P(\lambda) / \int d\lambda \Omega P(\lambda)$):

$$\begin{array}{lll} \langle \widehat{a}_i \rangle &=& E\left[\alpha_i\right] \\ \left\langle \, \widehat{a}_i^{\dagger} \widehat{a}_j \, \right\rangle &=& E\left[\alpha_i^{+} \alpha_j + n_{ij}\right] \\ \left\langle \, \widehat{a}_i \widehat{a}_j \, \right\rangle &=& E\left[\alpha_i \alpha_j + m_{ij}\right] \end{array}$$

 higher-order moments (normally ordered) are averages of Wick decompositions, eg:

$$\langle : \widehat{n}_i \widehat{n}_j : \rangle = E \left[n_{ii} n_{jj} - n_{ij} n_{ji} + m_{ij}^+ m_{ji} \right]$$

Non-Hermitian Elements

- when $\alpha^* = \alpha^+$, $\mathbf{m}^{\dagger} = \mathbf{m}^+$, and $\mathbf{n}^{\dagger} = \mathbf{n}$, the Gaussian operators correspond to density operators for physical states
- full Gaussian basis also includes non-Hermitian generalisations of these operators
 - → gives the overcompleteness necessary to ensure properties
 - 3: positivity, and
 - 4: freedom to diffuse
- for more on the properties of generalised Gaussian operators, see: J. F. Corney and P. D. Drummond, J. Phys. A 39, 269 (2006).
- the Hubbard-Stratonovich transformation, as used in Auxiliary field methods, gives exponentials of one-body operators (and thus Gaussian operators), but the approach is somewhat different. See e.g. S. Roumbouts and K. Heyde, Phys. Status. Solidi B 237, 99 (2003).

Differential Identities

all one-body operators equate to first-order derivatives:



Special case I: Coherent-state projectors

- for bosons, we can restrict the basis to $\vec{\lambda} = (\Omega, \alpha, (\alpha^+)^*, 0, 0, 0)$:
- get the kernal for the +P representation

$$\widehat{\Lambda} \;\; = \;\; \Omega rac{ig| lpha ig> ig< (lpha^+)^* ig|}{ig< (lpha^+)^* ig| lpha ig>}$$

differential identities reduce to

$$\widehat{\mathbf{a}}\widehat{\boldsymbol{\Lambda}} = \boldsymbol{\alpha}\widehat{\boldsymbol{\Lambda}}; \qquad \widehat{\mathbf{a}}^{\dagger}\widehat{\boldsymbol{\Lambda}} = \left(\boldsymbol{\alpha}^{+} + \frac{\partial}{\partial\boldsymbol{\alpha}}\right)\widehat{\boldsymbol{\Lambda}}$$
$$\widehat{\boldsymbol{\Lambda}}\widehat{\mathbf{a}} = \left(\boldsymbol{\alpha} + \frac{\partial}{\partial\boldsymbol{\alpha}^{+}}\right)\widehat{\boldsymbol{\Lambda}}; \qquad \widehat{\boldsymbol{\Lambda}}\widehat{\mathbf{a}}^{\dagger} = \boldsymbol{\alpha}^{+}\widehat{\boldsymbol{\Lambda}}$$

Special case II: Thermal operators

• for a number-conserving system, can restrict basis to $\overrightarrow{\lambda} = (\Omega, 0, 0, n, 0, 0)$:

$$\widehat{\mathbf{\Lambda}} = |\mathbf{I} \pm \mathbf{n}|^{\mp 1} : \exp\left[\widehat{\mathbf{a}}\left(\mathbf{I} \mp \mathbf{I} - [\mathbf{I} \pm \mathbf{n}]^{-1}\right)\widehat{\mathbf{a}}^{\dagger}\right] :$$

- defined for bosons (upper sign) and fermions (lower sign)
- moments:

$$\left\langle \widehat{a}_{i}\widehat{a}_{j}\right\rangle = 0; \qquad \left\langle \widehat{a}_{i}^{\dagger}\widehat{a}_{j}\right\rangle = E\left[n_{ij}\right]$$
$$\left\langle \widehat{a}_{i}^{\dagger}\widehat{a}_{j}^{\dagger}\widehat{a}_{j}\widehat{a}_{i}\right\rangle = E\left[n_{ii}n_{jj}\pm n_{ij}n_{ji}\right]$$

Special case II: Thermal operator identities

$$\begin{aligned} \widehat{\mathbf{b}}^{\dagger T} \widehat{\mathbf{b}}^{T} \widehat{\Lambda} &= \mathbf{n} \widehat{\Lambda} + (\mathbf{I} - \mathbf{n}) \frac{\partial \widehat{\Lambda}}{\partial \mathbf{n}} \mathbf{n}, \\ \widehat{\Lambda} \widehat{\mathbf{b}}^{\dagger T} \widehat{\mathbf{b}}^{T} &= \mathbf{n} \widehat{\Lambda} + \mathbf{n} \frac{\partial \widehat{\Lambda}}{\partial \mathbf{n}} (\mathbf{I} - \mathbf{n}), \\ \widehat{\mathbf{b}}^{\dagger T} \widehat{\Lambda} \widehat{\mathbf{b}}^{T} &= (\mathbf{I} - \mathbf{n}) \widehat{\Lambda} + (\mathbf{I} - \mathbf{n}) \frac{\partial \widehat{\Lambda}}{\partial \mathbf{n}} (\mathbf{I} - \mathbf{n}), \\ \left(\widehat{\mathbf{b}} \widehat{\Lambda} \widehat{\mathbf{b}}^{\dagger} \right)^{T} &= \mathbf{n} \widehat{\Lambda} - \mathbf{n} \frac{\partial \widehat{\Lambda}}{\partial \mathbf{n}} \mathbf{n} \end{aligned}$$

- where $ilde{\mathbf{n}} = \mathbf{I} \pm \mathbf{n}$ is the hole correlations
- for two-body operators, apply two of these identities in succession

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Density operators for quantum evolution

1. Unitary dynamics: $\widehat{\rho}(t) = e^{-i\widehat{H}t/\hbar}\widehat{\rho}(0)e^{i\widehat{H}t/\hbar}$

•
$$\frac{\partial}{\partial t}\widehat{\rho} = -\frac{i}{\hbar}\left[\widehat{H},\widehat{\rho}\right]$$

2. Equilibrium state: $\widehat{\rho}_{un}(T) = e^{-(\widehat{H} - \mu \widehat{N})/k_B T}$

•
$$\frac{\partial}{\partial\beta}\widehat{\rho} = \frac{1}{2}\left[\widehat{H} - \mu\widehat{N}, \widehat{\rho}\right]_{+}; \beta = 1/k_BT$$

3. Open dynamics: $\widehat{\rho}_{Sys} = Tr_{Res}\left\{\widehat{\rho}\right\}$

•
$$\frac{\partial}{\partial t}\widehat{\rho} = -\frac{i}{\hbar}\left[\widehat{H},\widehat{\rho}\right] + \gamma\left(2\widehat{R}\widehat{\rho}\widehat{R}^{\dagger} - \widehat{R}^{\dagger}\widehat{R}\widehat{\rho} - \widehat{\rho}\widehat{R}^{\dagger}\widehat{R}\right)$$

• each type is equivalent to a Liouville equation for $\widehat{\rho}$:

$$\frac{d}{d\tau}\widehat{\rho} = \widehat{L}[\widehat{\rho}]; \ \tau = t,\beta$$

Phase-space Recipe

- 1. Formulate: $\partial \widehat{\rho} / \partial \tau = \widehat{L}[\widehat{\rho}]$
- 2. Expand: $\int \partial P / \partial \tau \widehat{\Lambda} d \overrightarrow{\lambda} = \int P \widehat{L} \left[\widehat{\Lambda} \right] d \overrightarrow{\lambda}$
- 3. **Transform**: the operators to differential form: $\widehat{L} |\widehat{\Lambda}| = \mathcal{L} \widehat{\Lambda}$
- 4. Integrate by parts: $\int P \bot \widehat{\Lambda} d \overrightarrow{\lambda} \Longrightarrow \int \widehat{\Lambda} \bot' P d \overrightarrow{\lambda}$
 - neglect boundary terms (for sufficiently bounded P)
 - → analyse stability of resulting phase-space trajectories
 - \implies monitor tails of *P* during simulation
- 5. **Obtain** Fokker-Planck equation: $\partial P/\partial \tau = \angle P$

Fokker-Planck equations

resulting Fokker-Planck equation is of the form

$$\frac{d}{dt}P(\overrightarrow{\lambda},\tau) = \left[-\sum_{a=0}^{p}\frac{\partial}{\partial\lambda_{a}}A_{a}(\overrightarrow{\lambda}) + \frac{1}{2}\sum_{a,b=0}^{p}\frac{\partial}{\partial\lambda_{a}}\frac{\partial}{\partial\lambda_{b}}D_{ab}(\overrightarrow{\lambda})\right]P(\overrightarrow{\lambda},\tau)$$

- A_a gives the drift (deterministic component)
- D_{ab} gives the diffusion (stochastic component)
- D_{ab} must be positive-definite (in terms of real variables)
 - always possible by appropriate choice of derivatives:

$$\partial/\partial\lambda_a=\partial/\partial\lambda_a^x=-i\partial/\partial\lambda_a^y$$

 freedom in derivatives comes from analyticity (overcompleteness)

Itô stochastic equations

 $\bullet\,$ sample with Itô stochastic equations for $\overrightarrow{\lambda}$

$$d\lambda_a(\tau) = A_a(\overrightarrow{\lambda}) d\tau + \sum_b B_{ab}(\overrightarrow{\lambda}) dW_b(\tau)$$

• where $dW_b(\tau)$ are Weiner increments, obeying

 $\langle dW_b(\tau) dW_{b'}(\tau') \rangle = \delta_{b,b'} \delta(\tau - \tau') d\tau$

- i. e. Gaussian white noise
- noise matrix B_{ab} is obtained by

$$D_{ab} = \sum_{c} B_{ac} B_{bc}$$

Stratonovich stochastic equations

- multiplicative SDEs do not obey the normal rules of calculus
 - must take care to choose the appropriate numerical algorithm
- solution to Itô equations is obtained by an explicit (Euler) algorithm
- for solution by a semi-implicit method, must start from equivalent Stratonovich equations:

$$d\lambda_a = \left[A_a(\overrightarrow{\lambda}) - \frac{1}{2}\sum_{bc} B_{cb}(\overrightarrow{\lambda}) \frac{\partial B_{ab}(\overrightarrow{\lambda})}{\partial \lambda_c}\right] d\tau + \sum_b B_{ab}(\overrightarrow{\lambda}) dW_b$$

i.e. the drift is modified

Stochastic Gauges

- mapping from Hilbert space to phase space not unique
 - → many "gauge" choices
- can alter noise terms B_{ij} (diffusion gauge¹)
- can introduce arbitrary drift functions $g_j(\vec{\lambda})$ (drift gauge²)

Weight $d\Omega = \Omega \left[U d\tau + g_j dW_j \right]$ Trajectory $d\lambda_i = A_i d\tau + B_{ij} [dW_j - g_j d\tau]$

- can also choose different bases, identities
- 1. L. I. Plimak, M. K. Olsen and M. J. Collett, Phys. Rev. A 64, 025801 (2001)
- P. Deuar and P. D. Drummond, Comp. Phys. Commun. 142, 442 (2001); Phys. Rev. A 66, 033812 (2002); J. Phys. A: Math. Gen. 39, 2723 (2006)

Interacting many-body physics



- many-body problems map to nonlinear stochastic equations
- calculations can be from first principles
- precision limited only by sampling error
- choose basis to suit the problem
- for more details on the Gaussian method for fermions, see J. F. Corney and P. D. Drummond, Phys. Rev. B 73 125112 (2006).

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Application: Hubbard model



$$\widehat{H} = -\sum_{ij,\sigma} t_{ij} \widehat{c}_{i,\sigma}^{\dagger} \widehat{c}_{j,\sigma} + U \sum_{j} \widehat{c}_{j,\uparrow}^{\dagger} \widehat{c}_{j,\downarrow}^{\dagger} \widehat{c}_{j,\downarrow} \widehat{c}_{j,\uparrow}$$

simplest model of an interacting Fermi gas on a lattice

- weak-coupling limit \rightarrow BCS transitions
- solid-state models; relevance to High- T_c superconductors

Solving the Hubbard Model

- only the 1D model is exactly solvable (Lieb & Wu, 1968)
- even then, not all correlations can be calculated
- higher dimensions can use Quantum Monte Carlo methods.
- ★ except for a few special symmetrical cases, QMC suffers from sign problems with the Hubbard model
 - e.g. sign problems for repulsive interaction away from half filling
- sign problem increases with dimension, lattice size, interaction strength

Applying the Gaussian representation

Use thermal basis, and apply mappings

$$\begin{split} \widehat{\mathbf{n}}_{\sigma} \widehat{\boldsymbol{\rho}} & \to & \left\{ 2\mathbf{n}_{\sigma} - (\mathbf{I} - \mathbf{n}_{\sigma}) \frac{\partial}{\partial \mathbf{n}_{\sigma}} \mathbf{n}_{\sigma} \right\} P(\Omega, \mathbf{n}_{\uparrow}, \mathbf{n}_{\downarrow}) \\ \widehat{\boldsymbol{\rho}} \widehat{\mathbf{n}}_{\sigma} & \to & \left\{ 2\mathbf{n}_{\sigma} - \mathbf{I} - \mathbf{n}_{\sigma} \frac{\partial}{\partial \mathbf{n}_{\sigma}} (\mathbf{I} - \mathbf{n}_{\sigma}) \right\} P(\Omega, \mathbf{n}_{\uparrow}, \mathbf{n}_{\downarrow}) \\ \widehat{\boldsymbol{\rho}} & \to & - \frac{\partial}{\partial \Omega} \Omega P(\Omega, \mathbf{n}_{\uparrow}, \mathbf{n}_{\downarrow}) \end{split}$$

- \implies Fokker-Planck equation for P, with drift and diffusion
- \implies sample with stochastic equations for Ω and \mathbf{n}_{σ}

Positive-Definite Diffusion

Modify interaction term with a 'Fermi gauge':

$$U\sum_{j}:\widehat{n}_{jj,\downarrow}\widehat{n}_{jj,\uparrow}: = -rac{1}{2}|U|\sum_{j}:\left(\widehat{n}_{jj,\downarrow}-rac{U}{|U|}\widehat{n}_{jj,\uparrow}
ight)^{2}:$$

 \implies diffusion matrix has a real 'square root' matrix

- \implies realise the diffusion with a real noise process
- \implies problem maps to a real (and much more stable) subspace
- \implies weights Ω guaranteed to be *positive*

Itô Equations

Itô stochastic equations, in matrix form:

$$\frac{d\mathbf{n}_{\sigma}}{d\tau} = -\frac{1}{2} \left\{ (\mathbf{I} - \mathbf{n}_{\sigma}) \mathbf{T}_{\sigma}^{(1)} \mathbf{n}_{\sigma} + \mathbf{n}_{\sigma} \mathbf{T}_{\sigma}^{(2)} (\mathbf{I} - \mathbf{n}_{\sigma}) \right\}$$

$$\frac{d\Omega}{d\tau} = \Omega \left\{ \sum_{ij,\sigma} t_{ij} n_{ij,\sigma} - U \sum_{j} n_{jj,\downarrow} n_{jj,\uparrow} + \mu \sum_{j,\sigma} n_{jj,\sigma} \right\}$$

where the stochastic propagator matrix is

$$T_{ij,\sigma}^{(r)} = -t_{ij} + \delta_{ij} \left\{ U n_{jj,\sigma'} - \mu \pm \sqrt{2|U|} \xi_j^{(r)} \right\}$$

• $\xi_{i}^{(r)}$ are delta-correlated white noises:

$$\left\langle \xi_{j}^{(r)}(\tau) \, \xi_{j'}^{(r')}(\tau') \right\rangle = \delta_{j,j'} \delta_{r,r'} \delta(\tau - \tau')$$

Stratonovich Equations

in Stratonovich form, the stochastic propagator matrix is

$$T_{ij,\sigma}^{(r)} = -t_{ij} + \delta_{ij} \left\{ U\left(n_{jj,\sigma'} - n_{jj,\sigma} + \frac{1}{2}\right) - \mu \pm \sqrt{2|U|}\xi_j^{(r)} \right\}$$

- we use an iterative, semi-implicit algorithm¹, with an adaptive step-size to overcome stiffness
- for more on the Hubbard calculation, see 2 and 3
- 1. P.D. Drummond and I.K. Mortimer, J. Comp. Phys. 93 (1991) 144.
- 2. J. F. Corney and P. D. Drummond, Phys. Rev. Lett. 93, 260401 (2004)
- 3. P. D. Drummond and J. F. Corney, Comput. Phys. Commun. 169, 412 (2005)

1D Lattice-100 sites



Branching

averages are weighted,eg

$$ig\langle \widehat{\mathbf{n}}(\mathbf{\tau}) ig
angle \ = \ rac{\sum_{j=1}^{N_p} \mathbf{\Omega}^{(j)}(\mathbf{\tau}) \mathbf{n}^{(j)}(\mathbf{\tau})}{\sum_{j=1}^{N_p} \mathbf{\Omega}^{(j)}(\mathbf{\tau})}$$

✗ but weights spread exponentially → many irrelevant paths

 \rightarrow delete low-weight paths and clone high-weight paths:

$$m^{(jp)} = \text{Integer}\left[\xi + \Omega^{(jp)}/\overline{\Omega}\right]$$

- $\xi \in [0,1]$ is a random variable, $\overline{\Omega}$ is an average weight
- after branching, weights of surviving paths are equalised

16x16 2D Lattice



Symmetry Projection Schemes

- 4×4 , 16×16 calculations showed that sampling error was well controlled for various filling factors including those for which other methods suffer a severe sign problem
- more precise simulations at Würzburg and Zürich showed inaccuracies at large β for some correlation functions
- but true ground-state results could be obtained by supplementing with a symmetry projection scheme¹
 - trade-off is an increase in sampling error
- 1. F. F. Assaad, P. Werner, P. Corboz, E. Gull and M. Troyer, Phys. Rev. B 72, 224518 (2005)

Symmetry Preservation

- problem: chosen basis does not reflect all symmetries of the Hamiltonian
 - in principle, distribution will restore the symmetry
 - but broad tails mean that
 - accurate sampling is difficult
 - possibility of boundary-term errors
- possible solution:
 - use basis that possesses Hamiltonian symmetries
 - use stochastic gauges to ensure boundedness of P

Summary

- Generalised phase-space representations provide a means of simulating many-body quantum physics from first principles
- Coherent-state-based methods have been successful in simulating quantum dynamics of photons and weakly interacting bosons.
- Gaussian-based methods extend the applicability to highly correlated systems of bosons and *fermions*.
- Simulated the Hubbard model, apparently without sign errors.
- Some inaccuracies at low temperature can be overcome by:
 - supplementary symmetry projection, or
 - possibly by basis and stochastic gauge choice.

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